

CARTOGRAPHIC STUDY OF THE MEO PHASE SPACE FOR PASSIVE DEBRIS REMOVAL

Despoina K. Skoulidou, Aaron J. Rosengren, Kleomenis Tsiganis,
and George Voyatzis

dskoulid@physics.auth.gr, aaronjay@physics.auth.gr, tsiganis@auth.gr, voyatzis@auth.gr

Department of Physics, Aristotle University of Thessaloniki, 54 124 Thessaloniki, Greece



Introduction

Recently, it has been realized that lunisolar secular resonances are of particular importance in the MEO regime, which may provide a substantial perigee boost for the satellite under properly chosen circumstances [1, 2, 3, 4]. Additionally, solar radiation pressure (SRP) plays crucial role in the motion of satellites, especially at large altitudes [5, 6]. We use a suitably modified version of the SWIFT symplectic integration package [7], to integrate the trajectories of ~ 7.5 m initial conditions in the GNSS and GTO regions for 120 *yr*, assuming nominal and enhanced SRP (e.g. solar sail).

Model & Initial conditions

- Earth (J_2 , C_{22} , S_{22}) + Moon + Sun (no expansion)
- Nominal C_R (A/m) = $0.015 \text{ m}^2/kg$ (SRP₁) and enhanced C_R (A/m) = $1 \text{ m}^2/kg$ (SRP₂)
- Two initial epochs (in 2018 and 2020)
- $t_{max} = 120 \text{ yr}$ & Re-entry limit $h_R = 400 \text{ km}$
- code: SWIFT-SAT, symplectic integrator, non-averaged equations of motion

Used to build a complete dynamical map of circum-terrestrial space (check presentation by A. Rossi)

Maps of the MEO phase space

- $a \in [0.6 - 0.71]$ (a_{GEO}), $e \in [0 : 0.88]$, $i \in [0 : 90^\circ]$
- $\Delta\Omega$, $\Delta\omega = \{0, 90^\circ, 180^\circ, 270^\circ\}$ w.r.t. equatorial values of Moon

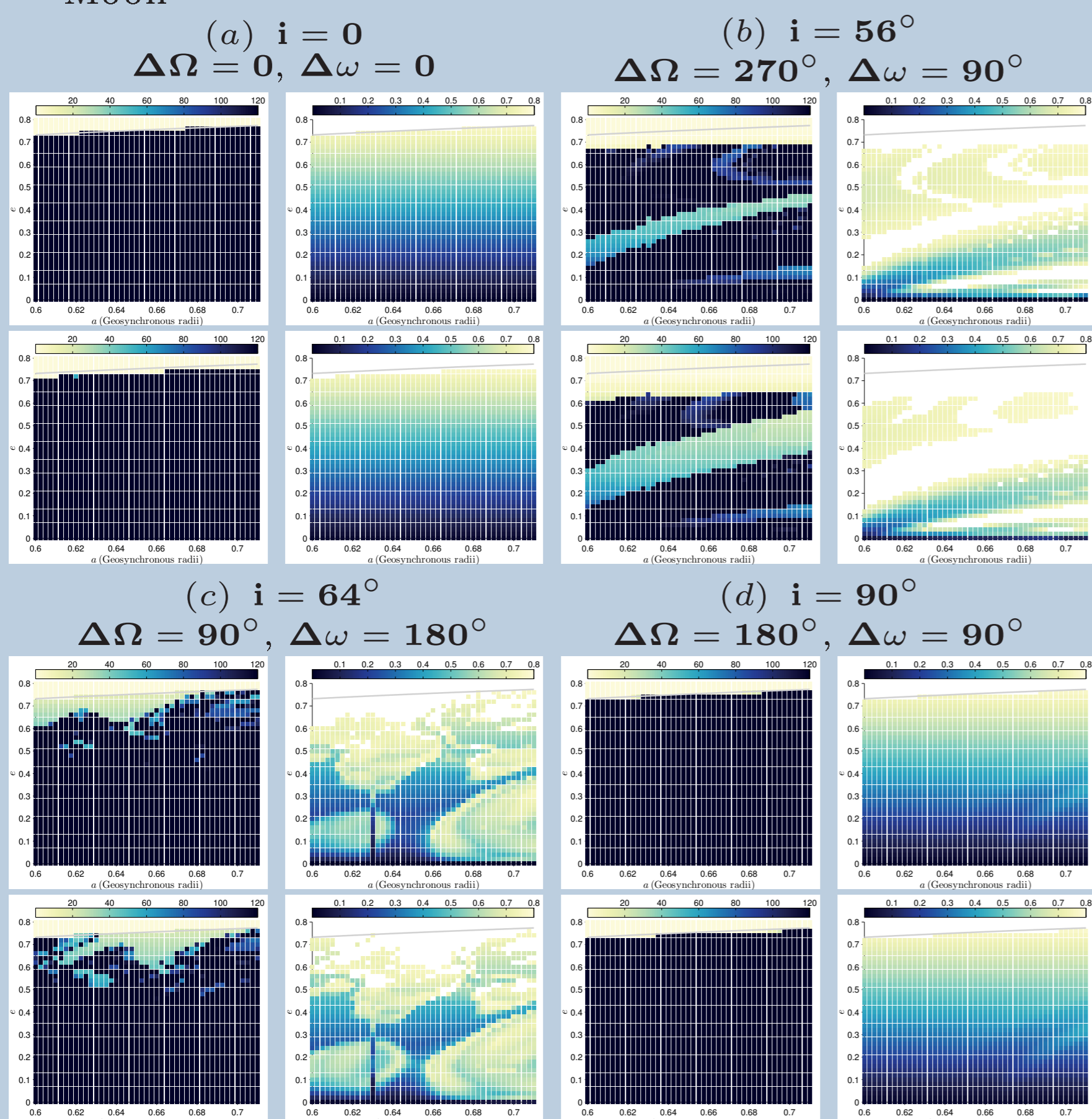


Figure 1: Lifetime (left) and sup-e maps (right) for SRP₁ (top) & SRP₂ (bottom).

- Re-entry solutions feasible, but not for $t < 25 \text{ yrs}$
- Enhanced A/m can widen the re-entry regions
- Results depend on secular orientation angles and epoch

Maps for GTO phase space

- $a \in [0.498 - 0.664]$ (a_{GEO}), $e \in [0.5 : 0.8]$
- $i = i_o \pm 10^\circ$, $i_o \sim 5^\circ$ (Kourou), $\sim 28^\circ$ (KSC), $\sim 46^\circ$ (Baikonur)
- $\Delta\Omega$, $\Delta\omega = \{0, 90^\circ, 180^\circ, 270^\circ\}$ w.r.t. ecliptic values of Moon

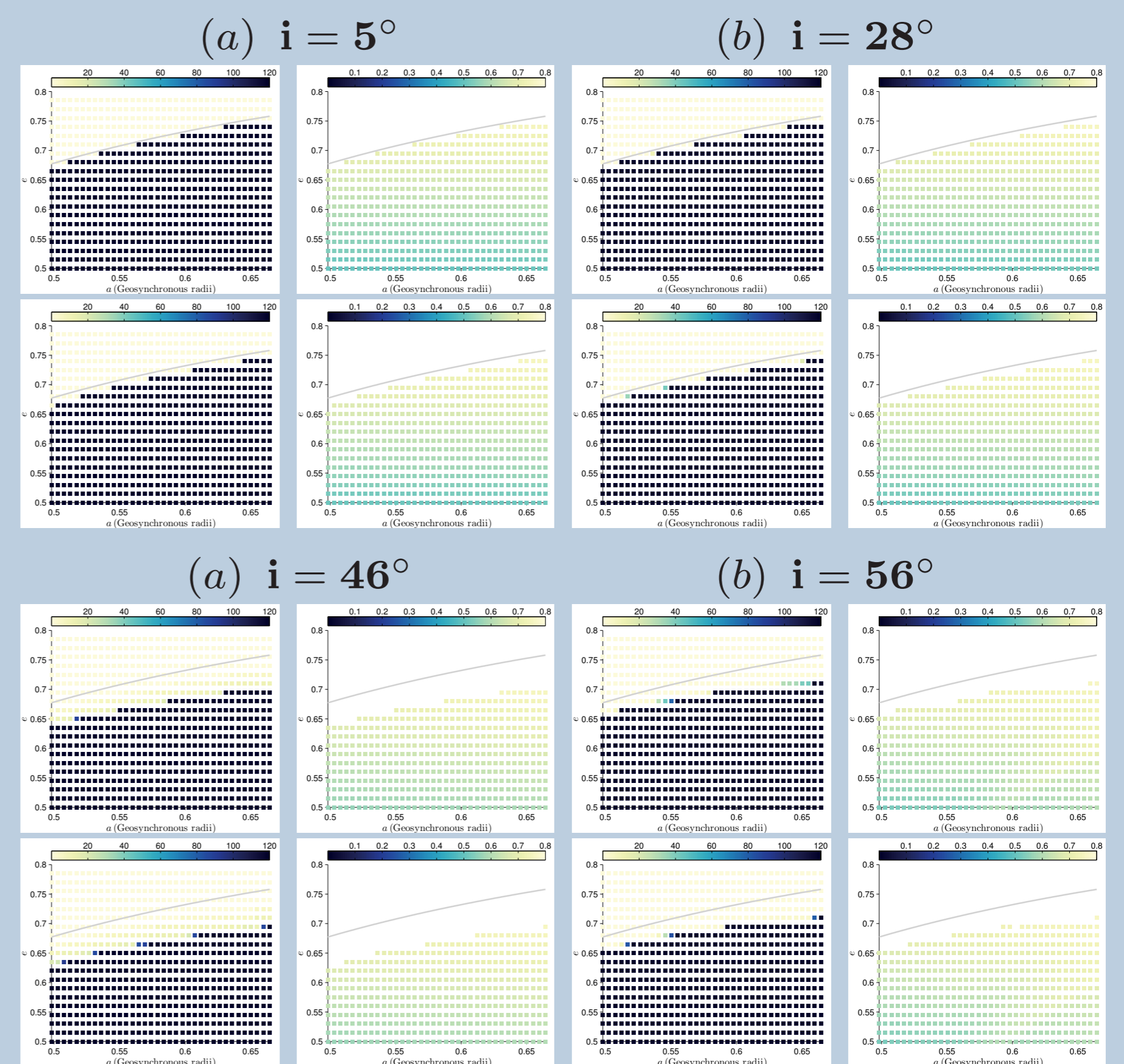


Figure 2: Lifetime (left) and sup-e maps (right) for $\Delta\Omega = 90^\circ$, $\Delta\omega = 0$ and SRP₁ (top) & SRP₂ (bottom).

- SRP is not very important, but helps
- Re-entry is easier at high inclinations

Conclusions - Future work

GNSS:

- Re-entry times of 40 – 60 *yr* feasible
- Design of maneuvers to reach the optimal diposal orbit for each initial operational orbit

GTO:

- Enhanced SPR is not enough for natural re-entry
- Atmospheric drag may help - need to check

References

1. Rosengren A. J., Alessi E. M., Rossi A., and Valsecchi G. B., *MNRAS* **449**, 3522, 2015
2. Daquin J., Rosengren A. J., Alessi E. M., Deleflie F., Valsecchi G. B., and Rossi A., *Celestial Mechanics and Dynamical Astronomy* **124**, 335, 2016
3. Celletti A., Gales C., and Pucacco G., *SIAM Journal on Applied Dynamical Systems* **15**, 1352, 2016
4. Gkolias I., Daquin J., Gachet F., and Rosengren A. J., *The Astronomical Journal* **152**, 119 (15pp), 2016
5. Colombo C., Lucking C., and McInnes C. R., *Acta Astronautica* **81**, 137, 2012
6. Lantukh D., Russell R. P., S. Broschart S., *Celestial Mechanics and Dynamical Astronomy* **121**, 171, 2015
7. Levison H., and Duncan M., *Icarus* **108**, 18, 1994

This research is funded by the European Commissions Horizon 2020, Grant Agreement 687500 (project ReDSHIFT).

<http://redshift-h2020.eu/>